Predicting Muon Fluxes and Seasonal Variations in Underground and Underwater Labs Using MUTE

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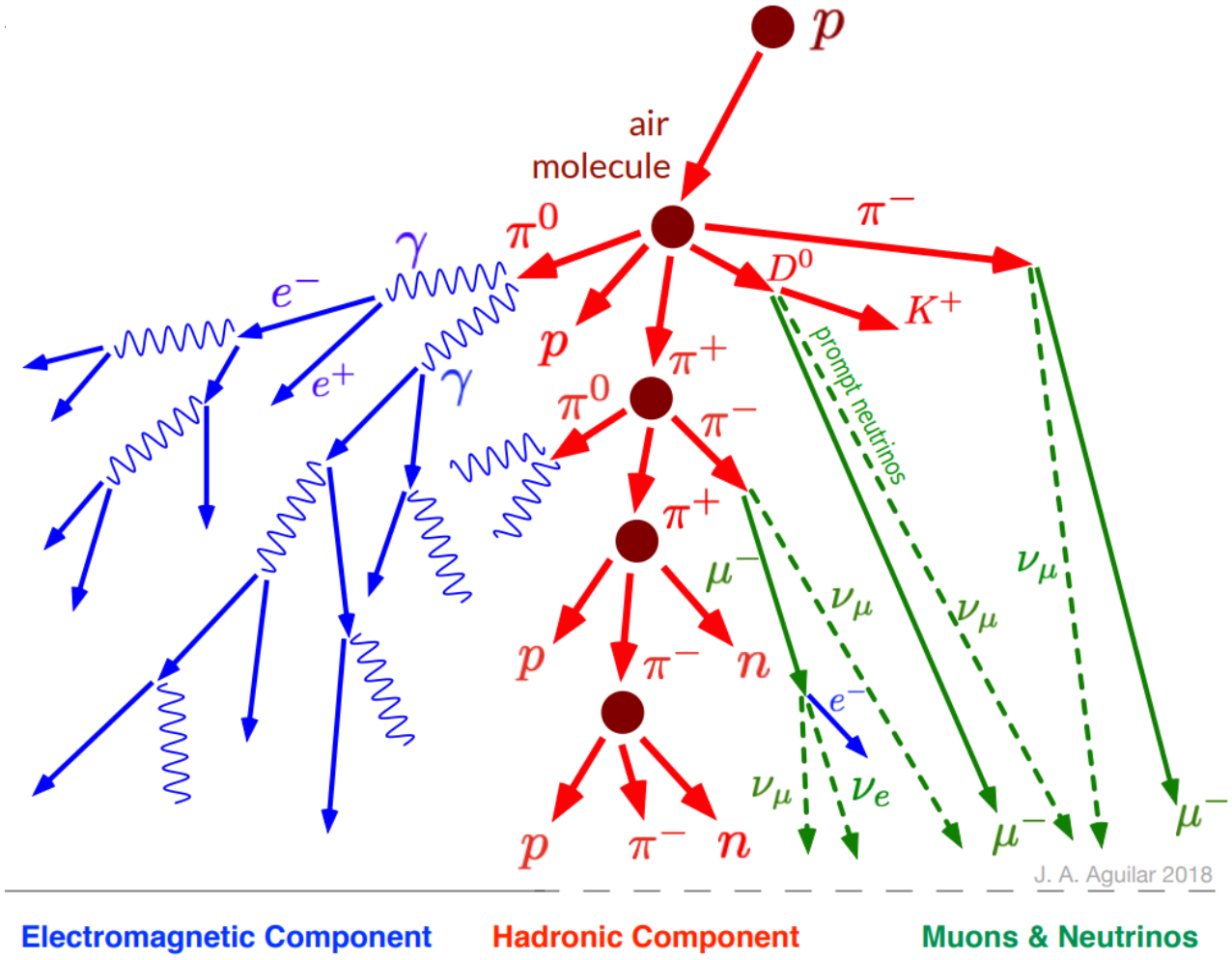
Introduction

- Cosmic rays interact in Earth's middle atmosphere to produce muons [1].
- Muons can easily penetrate matter by multiple kilometres.
- Underground and underwater muons are crucial in data analyses and the design of Dark Matter and neutrino detectors.
- Therefore, good knowledge of their flux is important in calculations of expected muon-induced backgrounds.



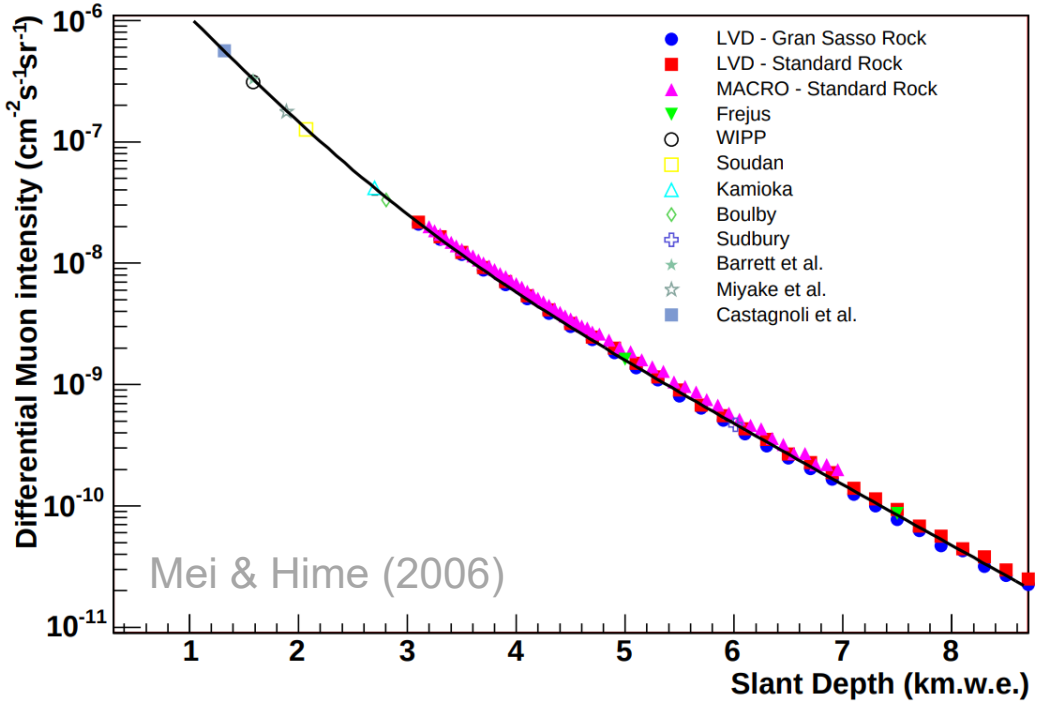
[1] J. Aguilar. *PHYS 467: Astroparticle Physics*. Lecture Notes.

Depth-Intensity Relations

- Depth Intensity Relations [2, 3] are a common way of calculating underground muon fluxes.
- Disadvantages:
  1. They are simple parametric fits.
  2. They are susceptible to statistical errors at deep slant depths.
  3. They are approximate and introduce systematic errors for θ > ~20° [4].
- MUTE (**MU**on in**T**ensity cod**E**) solves all three of these problems.
- It is a computational tool written in Python that calculates muon spectra underground.

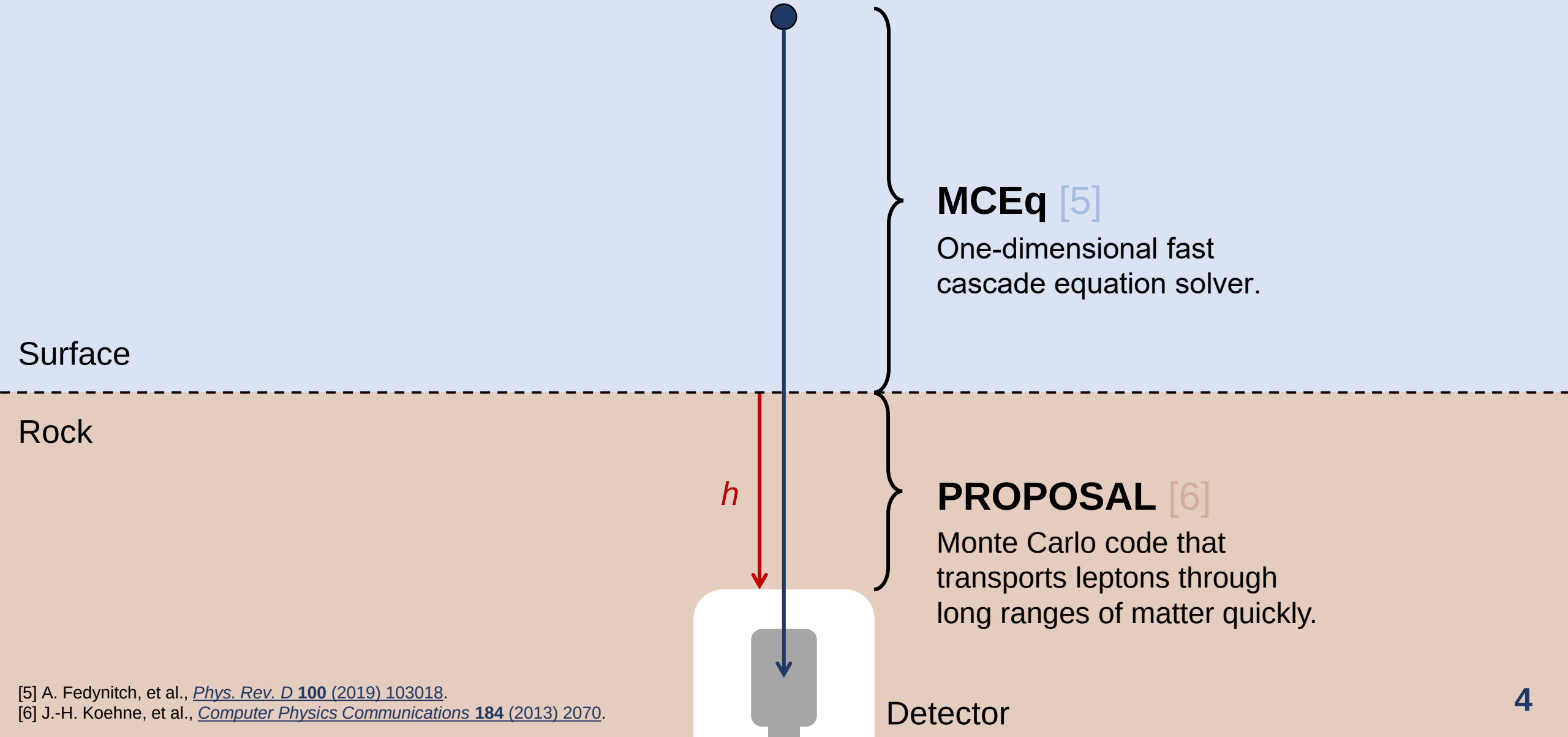
Continuous Losses **Discrete Losses**

$$I(h) = I_1 e^{(-h/\lambda_1)} + I_2 e^{(-h/\lambda_2)}$$



[2] D. Mei and A. Hime, *Phys. Rev. D* **73** (2006) 053004 [astro-ph/0512125].
[3] M. Crouch, in *ICRC*, vol. 6, p. 165, Jan., 1987.
[4] A. Fedynitch, W. Woodley and M.-C. Piro, *ApJ* **928** (2022) 27.

Method – Overview



MCEq [5]

One-dimensional fast cascade equation solver.

PROPOSAL [6]

Monte Carlo code that transports leptons through long ranges of matter quickly.

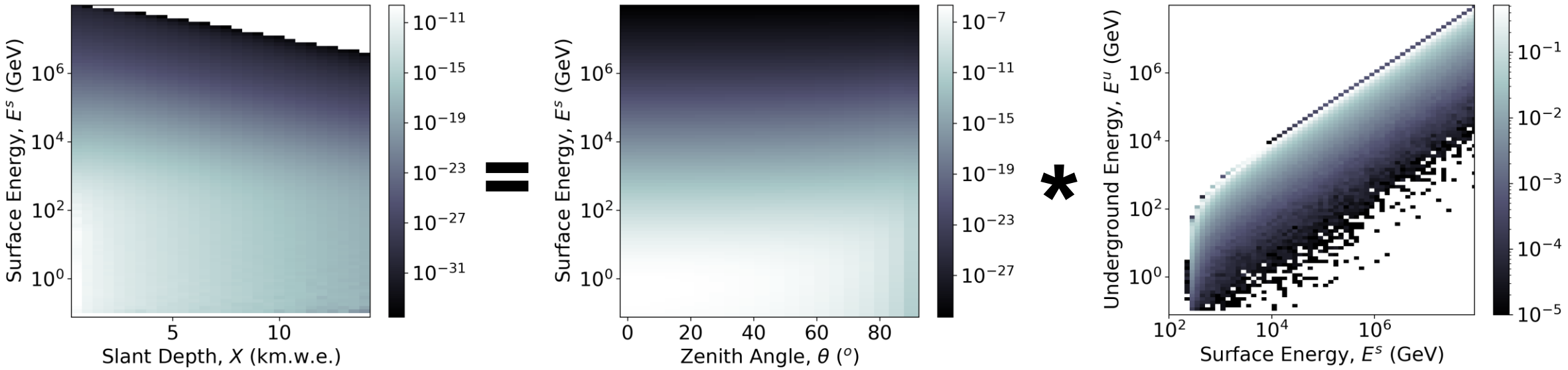
[5] A. Fedynitch, et al., *Phys. Rev. D* **100** (2019) 103018.
[6] J.-H. Koehne, et al., *Computer Physics Communications* **184** (2013) 2070.

Method – Convolution

- A convolution is performed to calculate underground fluxes:

$$\Phi^u(E_j^u, X_k, \theta_k) = \sum_i \Phi^s(E_i^s, \theta_k) P\left(E_i^s, E_j^u, X_k\right) \left(\frac{\Delta E_i^s}{\Delta E_j^u}\right)$$

Φ^u = Φ^s * P



MUTE
Underground Fluxes

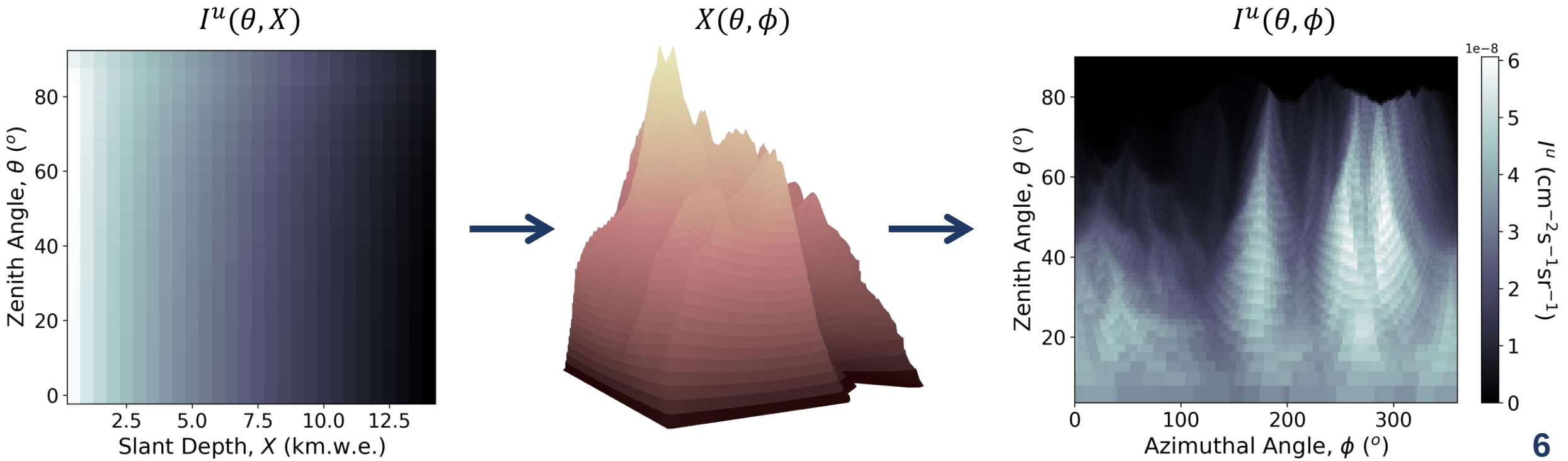
MCEq
Surface Fluxes

PROPOSAL
Transfer Tensor

Method – Labs under Mountains

- Underground intensities for mountains are first calculated on a grid of constant zenith angles and slant depths.

- Using a map of the mountain profile, these intensities are then interpolated to the slant depths $X(\theta, \phi)$ that define the mountain.



Method – Surface Flux Models

- The user can switch out models to be used in MCEq for surface flux calculations.

DDM [7]**:**

- Data-driven model for hadronic interactions.
- Uses low-energy accelerator data.
- Extrapolates to higher energies using Feynman scaling.

[7] A. Fedynitch and M. Huber, *Phys. Rev. D* **106** (2022) 083018 [2205.14766].
[8] J.P. Yañez and A. Fedynitch, 2303.00022.

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DAEMONFLUX [8]**:**

- Combines DDM and Global Spline Fit (GSF).
- Calibrated to muon flux and ratio data.
- Φ_ν uncertainties <10% up to 1 TeV.

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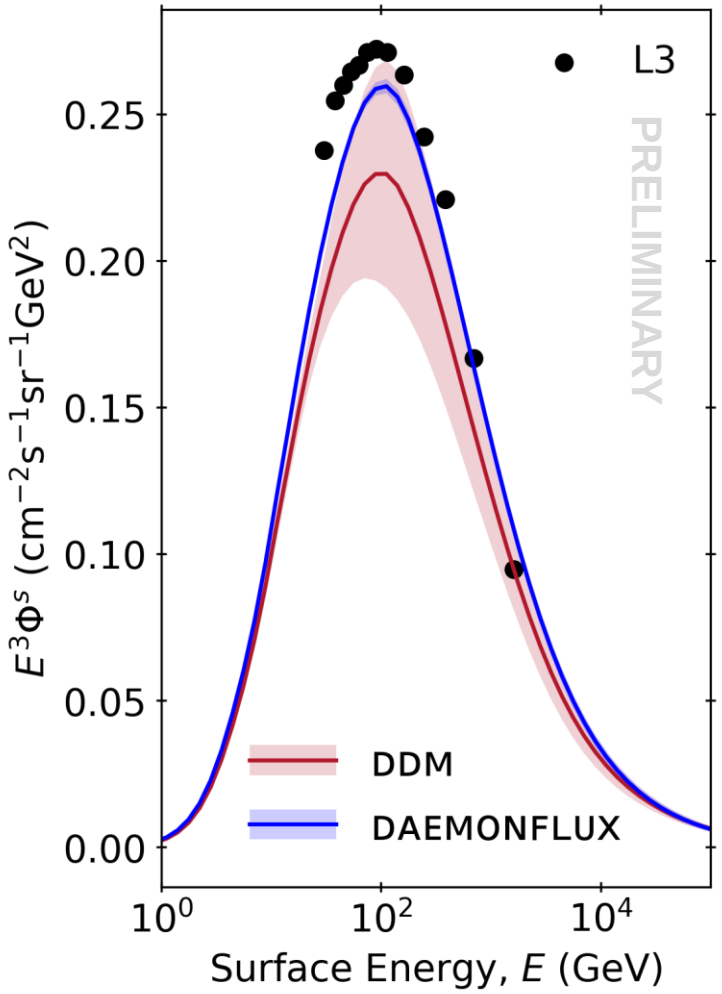

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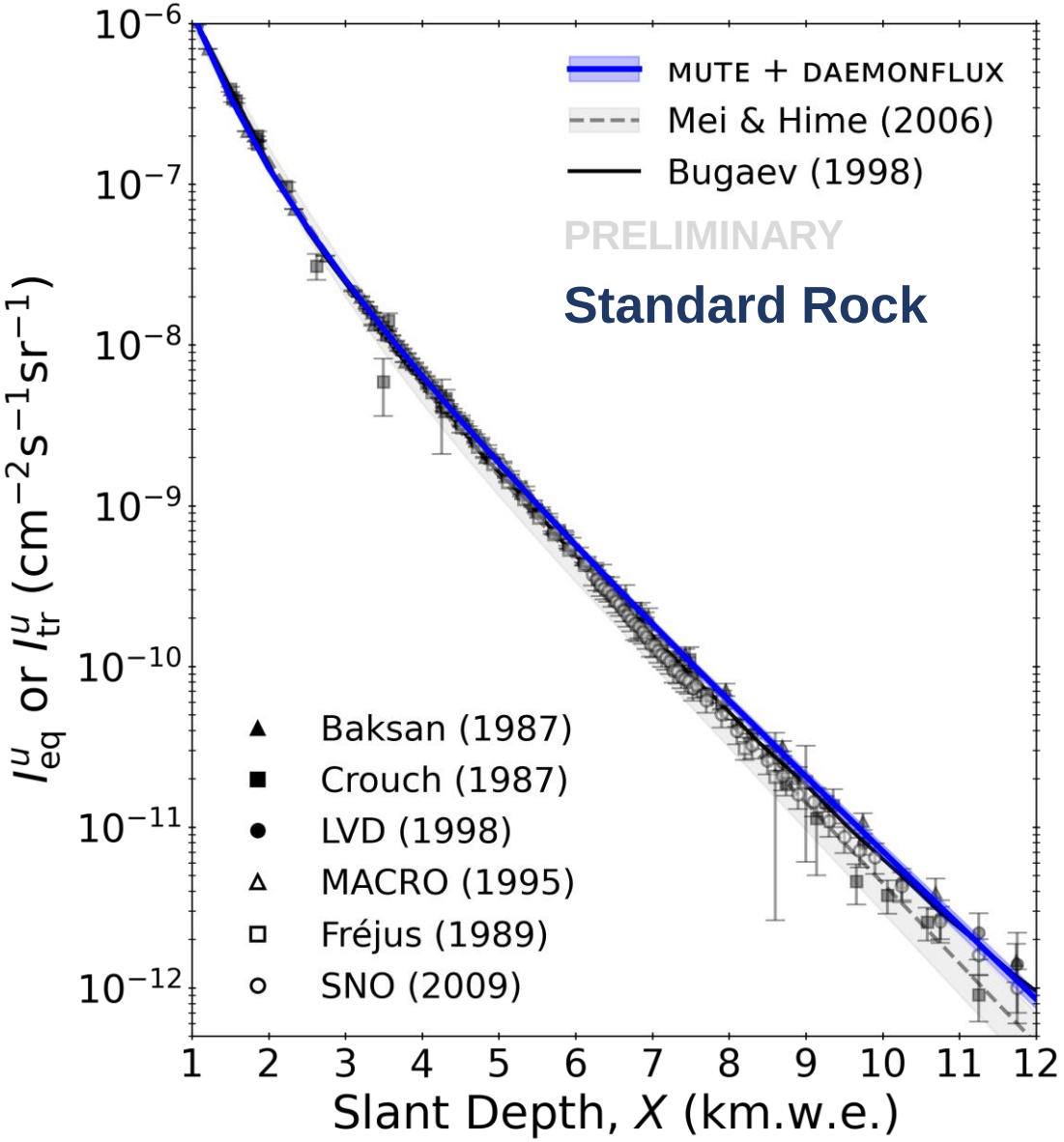
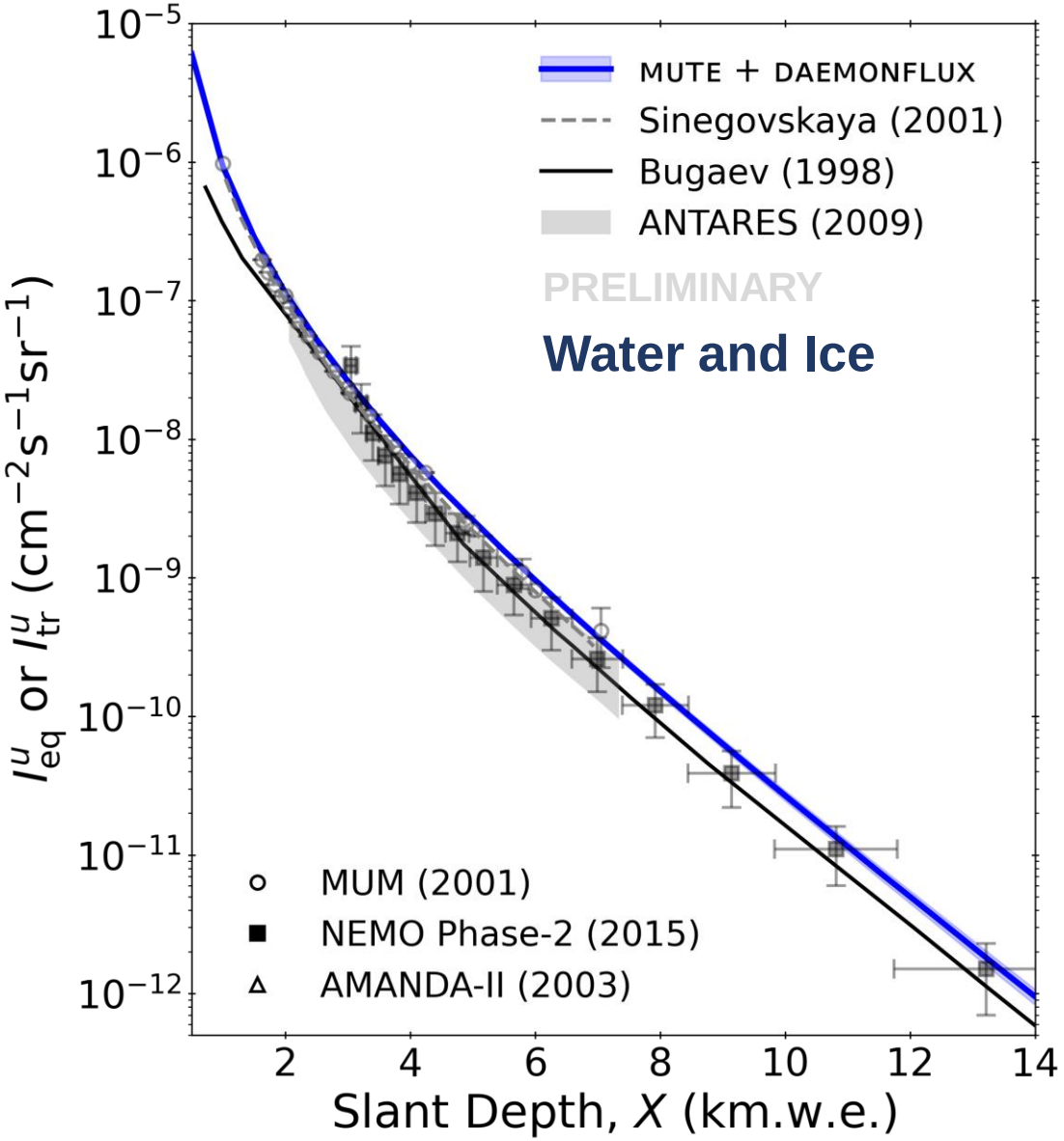
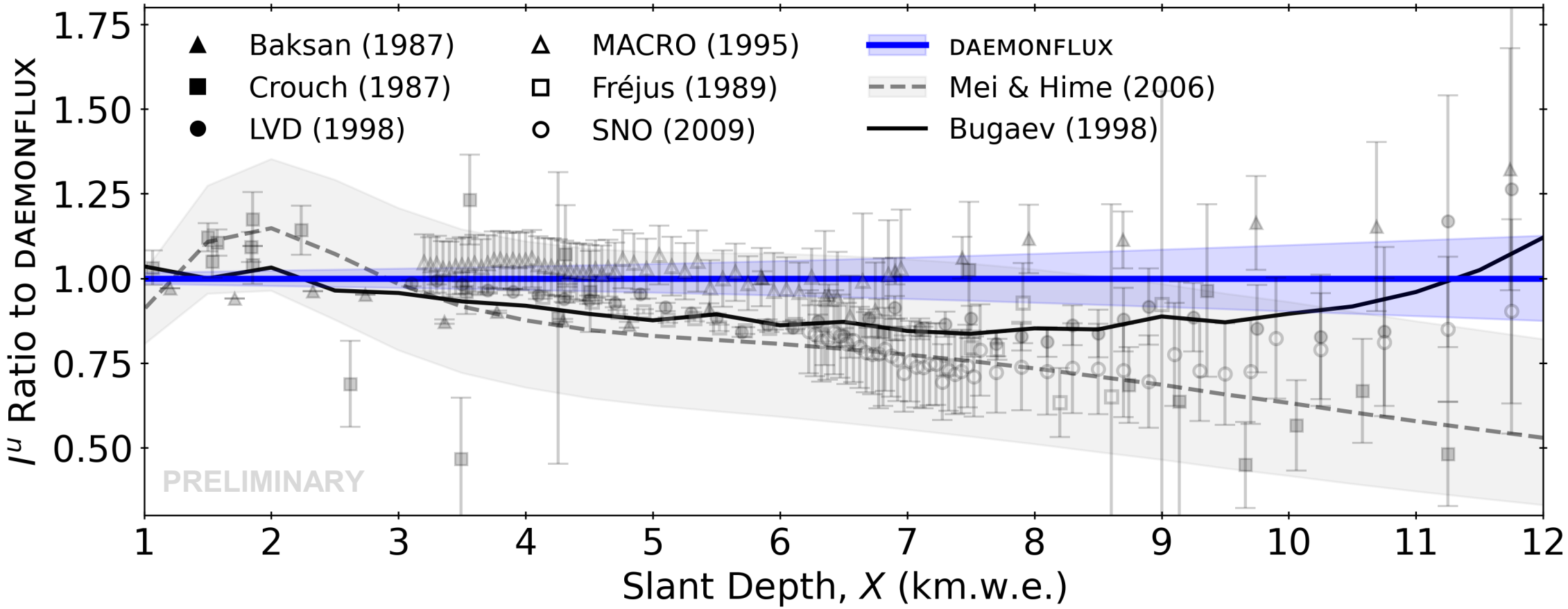
Results – Vertical Underground Intensity





Results – Comparison to Data

- Uncertainties have been reduced from 15% to ~1% by using DAEMONFLUX.



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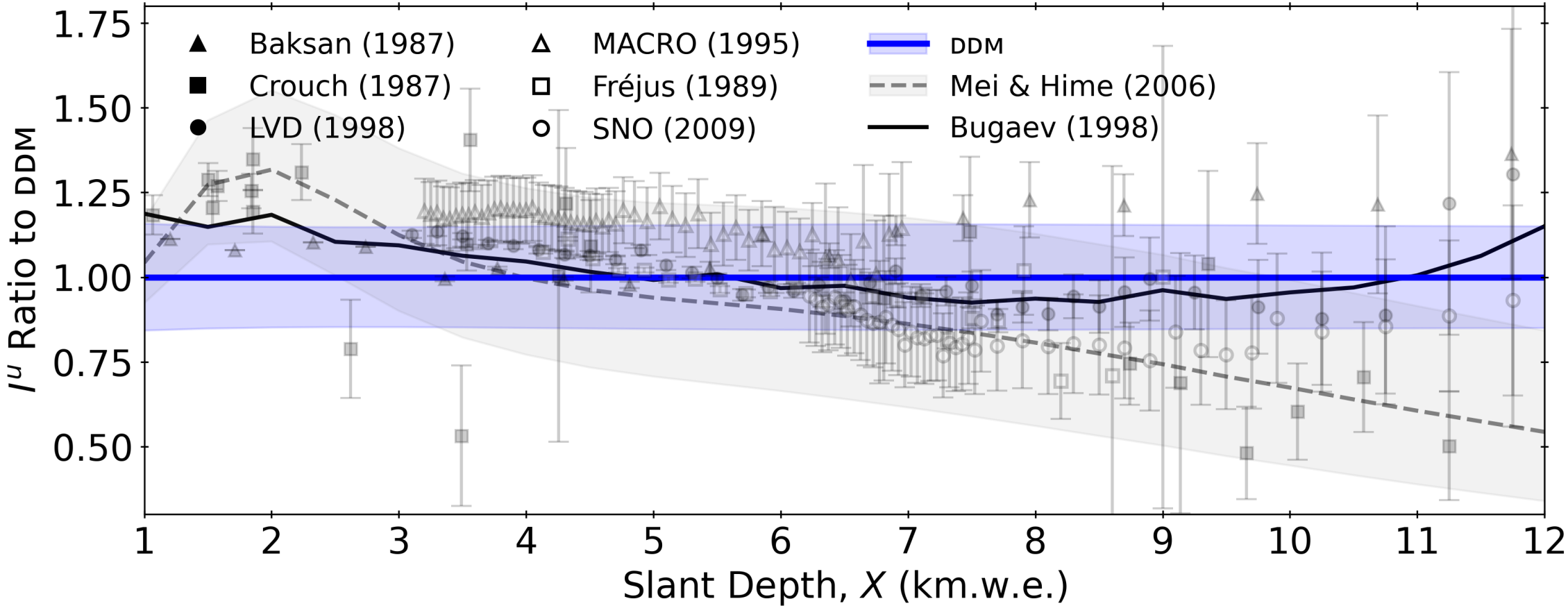
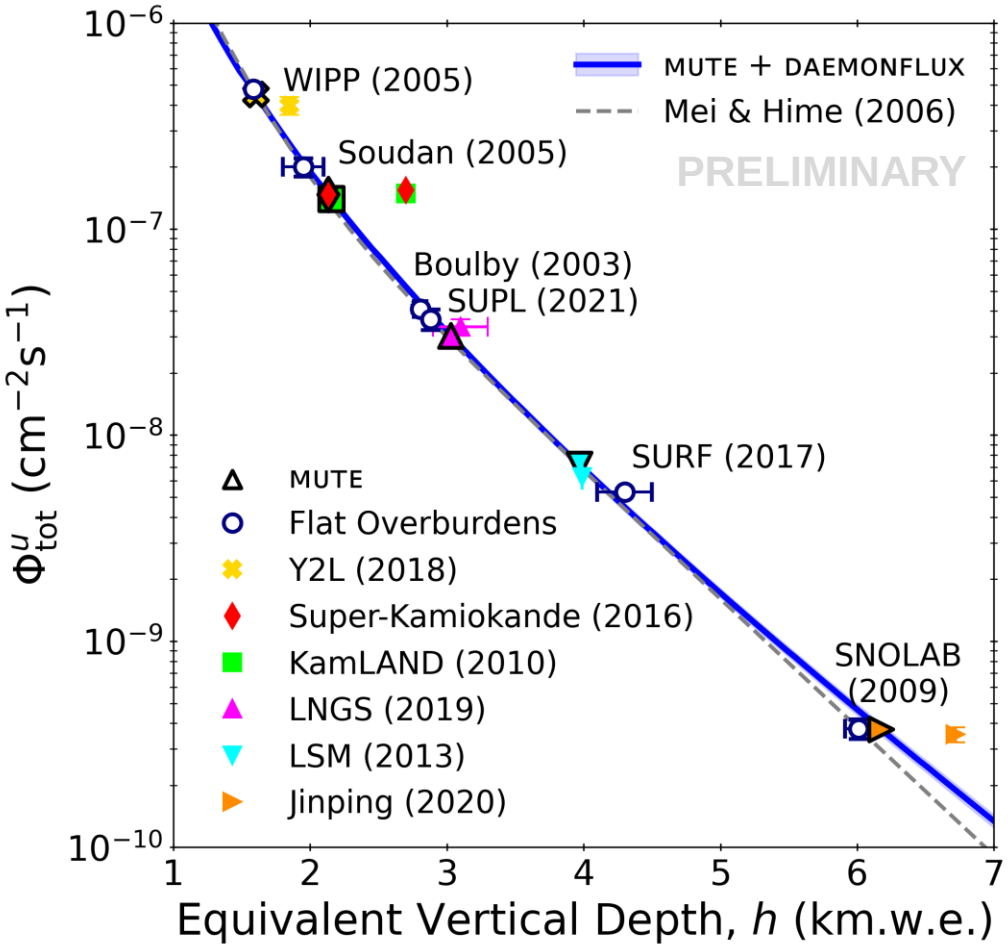
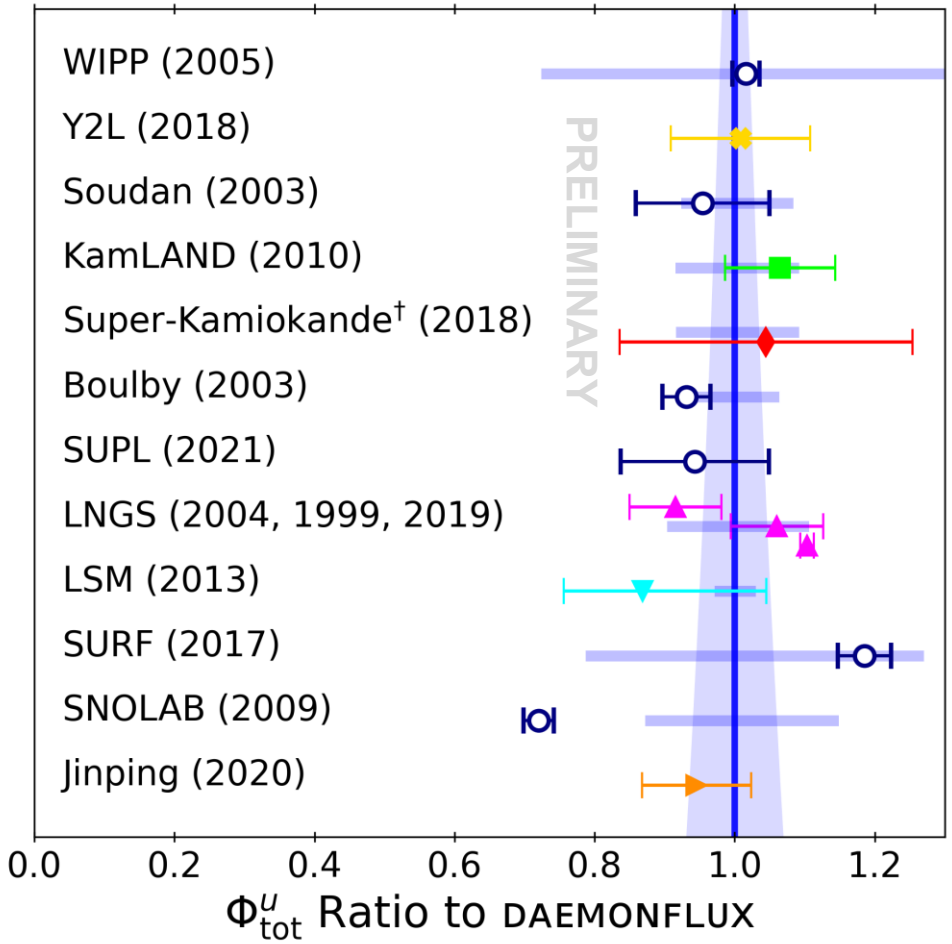


Results – Total Underground Flux

- Total flux calculations are consistent with measurements for labs under flat overburdens and mountains within theoretical errors in nearly every case.





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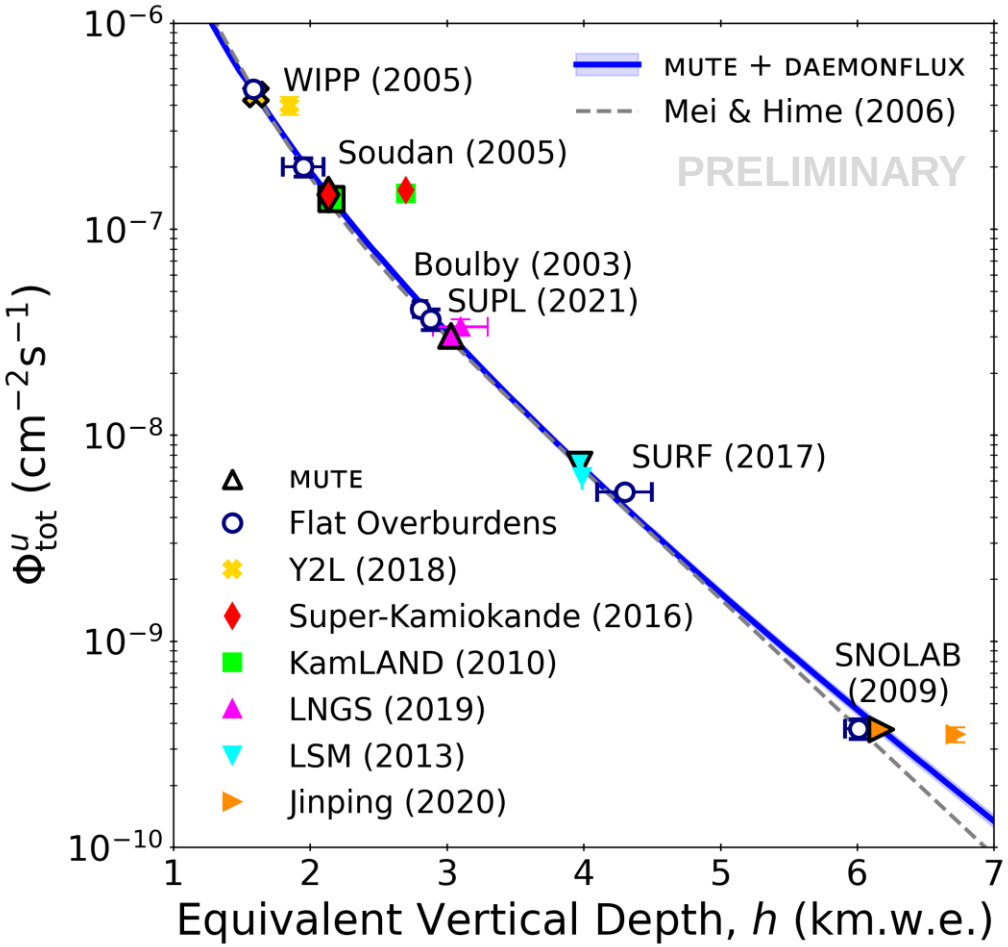

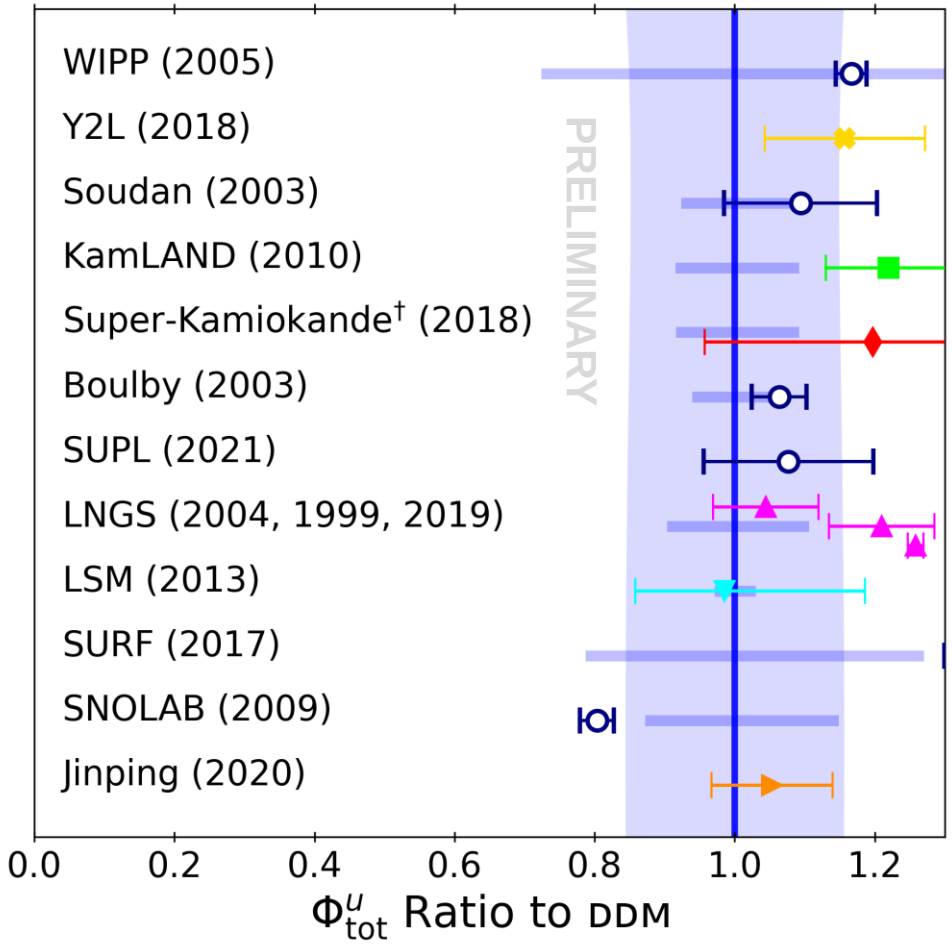

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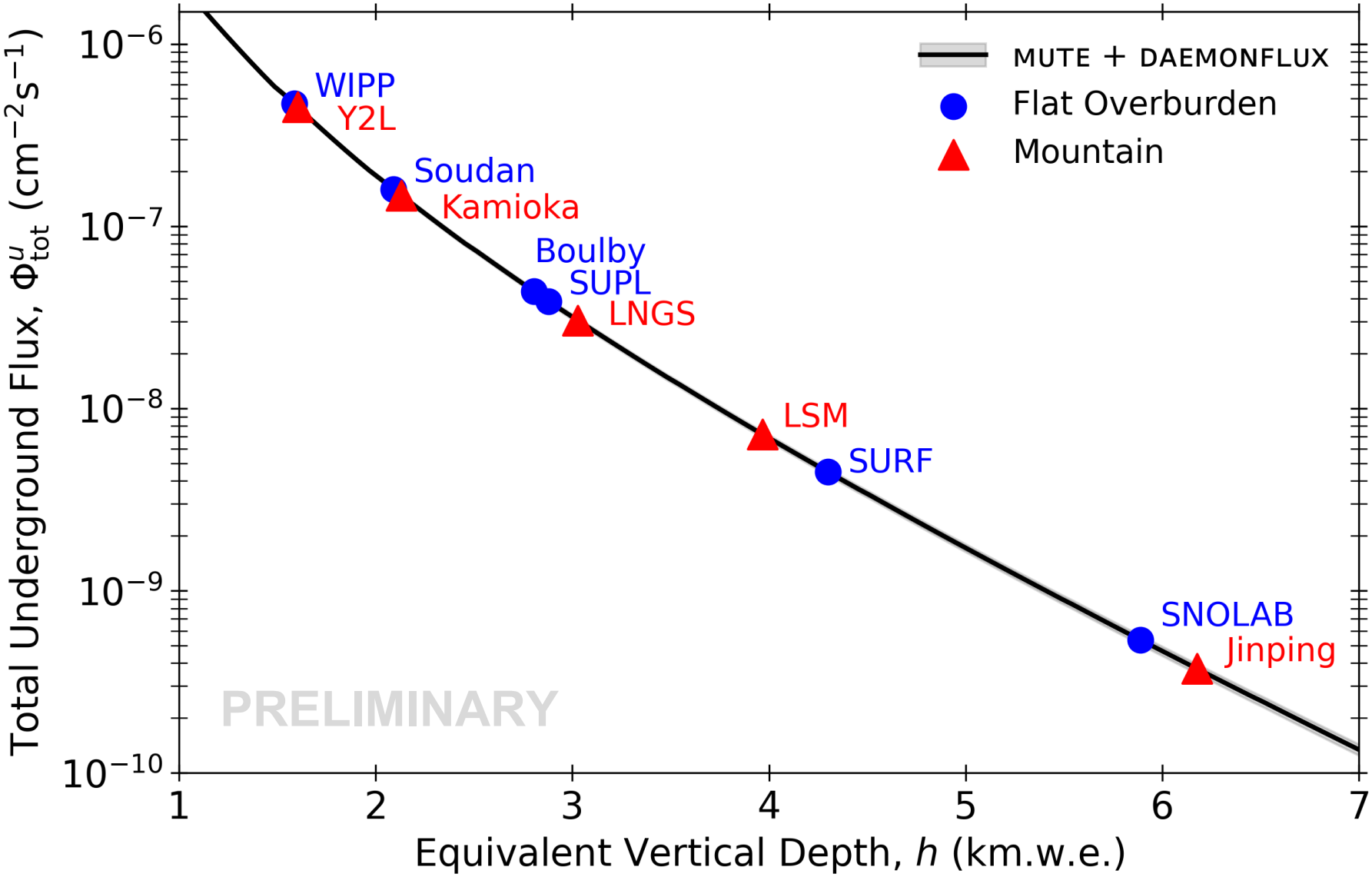
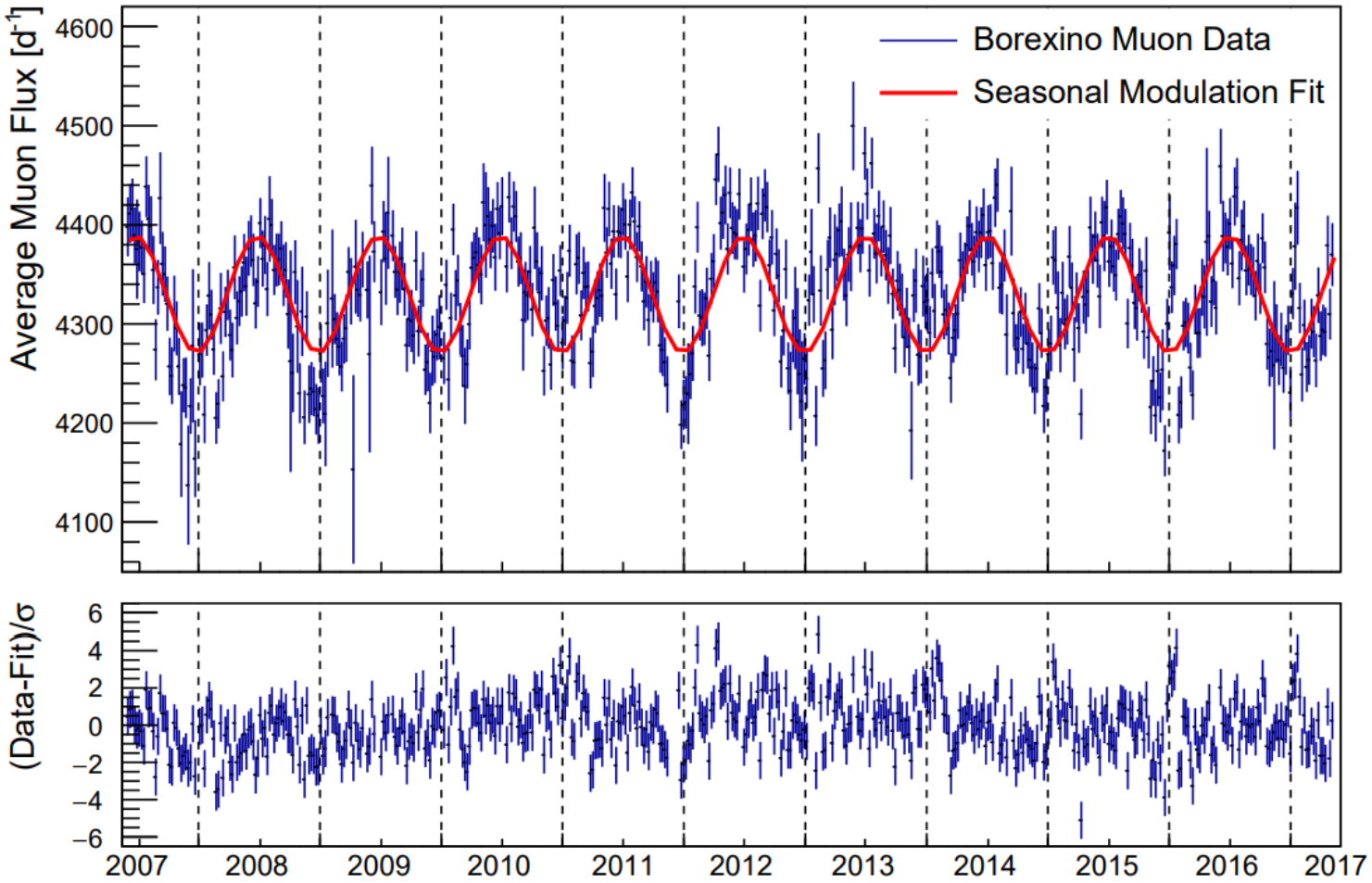


Seasonal Variations

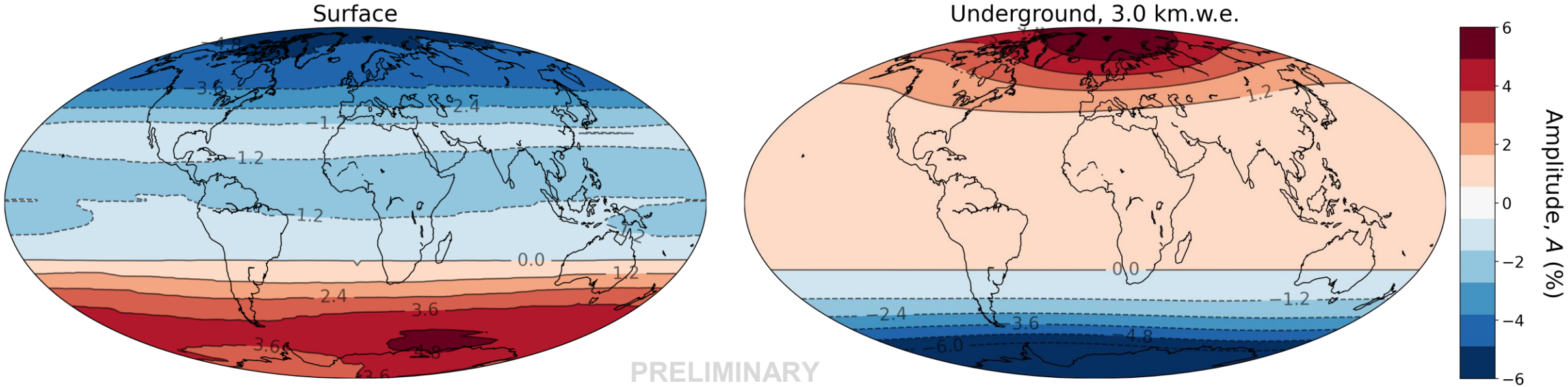
- The phenomenon of seasonal modulations in the muon flux is well-known [9]:



[9] M. Agostini, *et al*., *Journal of Cosmology and Astroparticle Physics* **2019** (2019) 046.

Seasonal Variations – Results

- I have calculated the amplitude of seasonal variations around the globe:

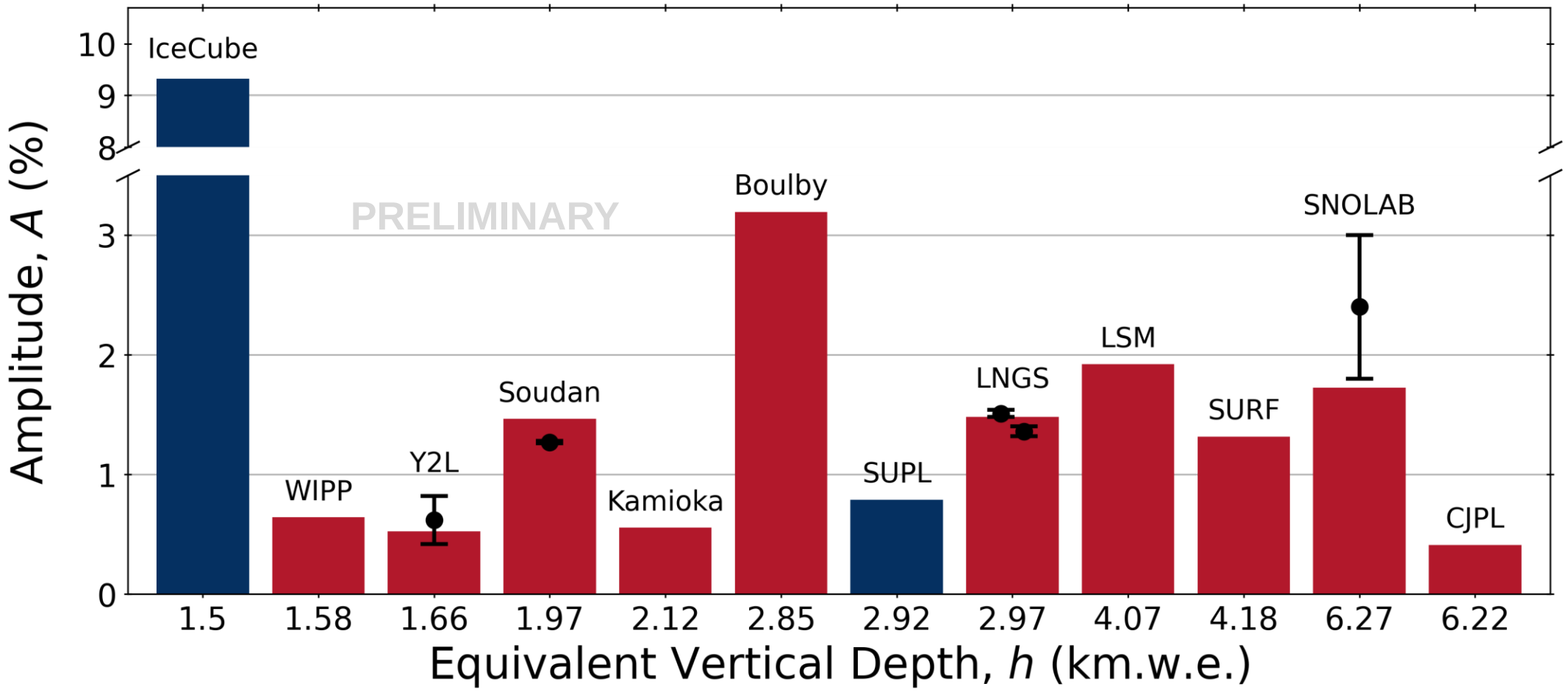


- The muon flux is **lower** at the surface in summer in the **northern hemisphere**.

- However, there are more higher-energy muons in the summer, which reach deeper underground. Therefore, the muon flux is **higher** underground in summer.

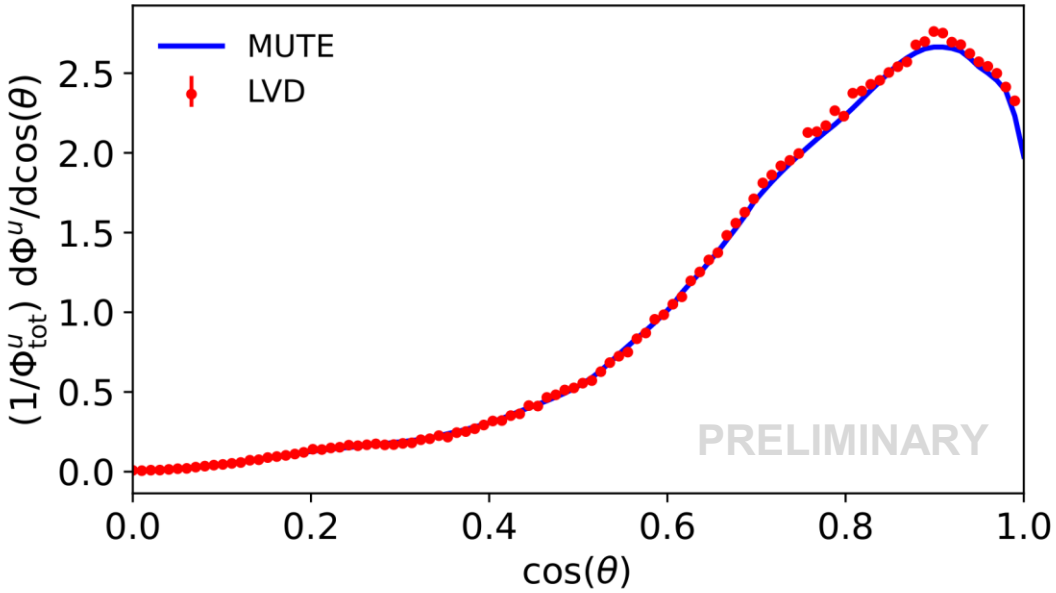
Seasonal Variations – Results

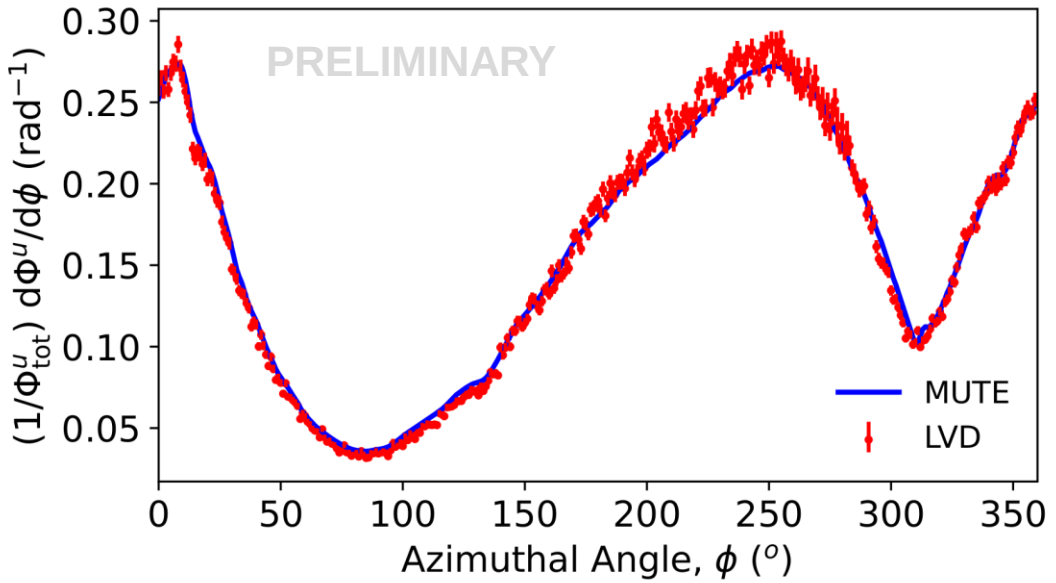
- MUTE can calculate seasonal variation amplitudes to high accuracy.



Applications – Angular Distributions

- MUTE can also calculate one-dimensional angular distributions for labs under mountains in the θ and ϕ directions.
- Results for the Gran Sasso mountain have been compared to data from the LVD experiment.
- We obtain very good agreement for the muon spectrum and flux, and for the shape of the mountain.
- This serves as a way of verifying the data analysis of the LVD experiment.





Conclusion

- MUTE is flexible, fast, and precise. It gives a full description of muon distributions underground and underwater, and can provide forward predictions for total muon fluxes.
- The results match experimental data very well for all physical observables. This can be used to cross-check data analyses.
- Uncertainties have been significantly reduced with the latest model, DAEMONFLUX.
- MUTE is public and available (`pip install mute`) to be used by experiments in labs under flat overburdens and mountains.

doi:10.3847/1538-4357/ac5027

https://github.com/wjwoodley/mute

Thank you

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